

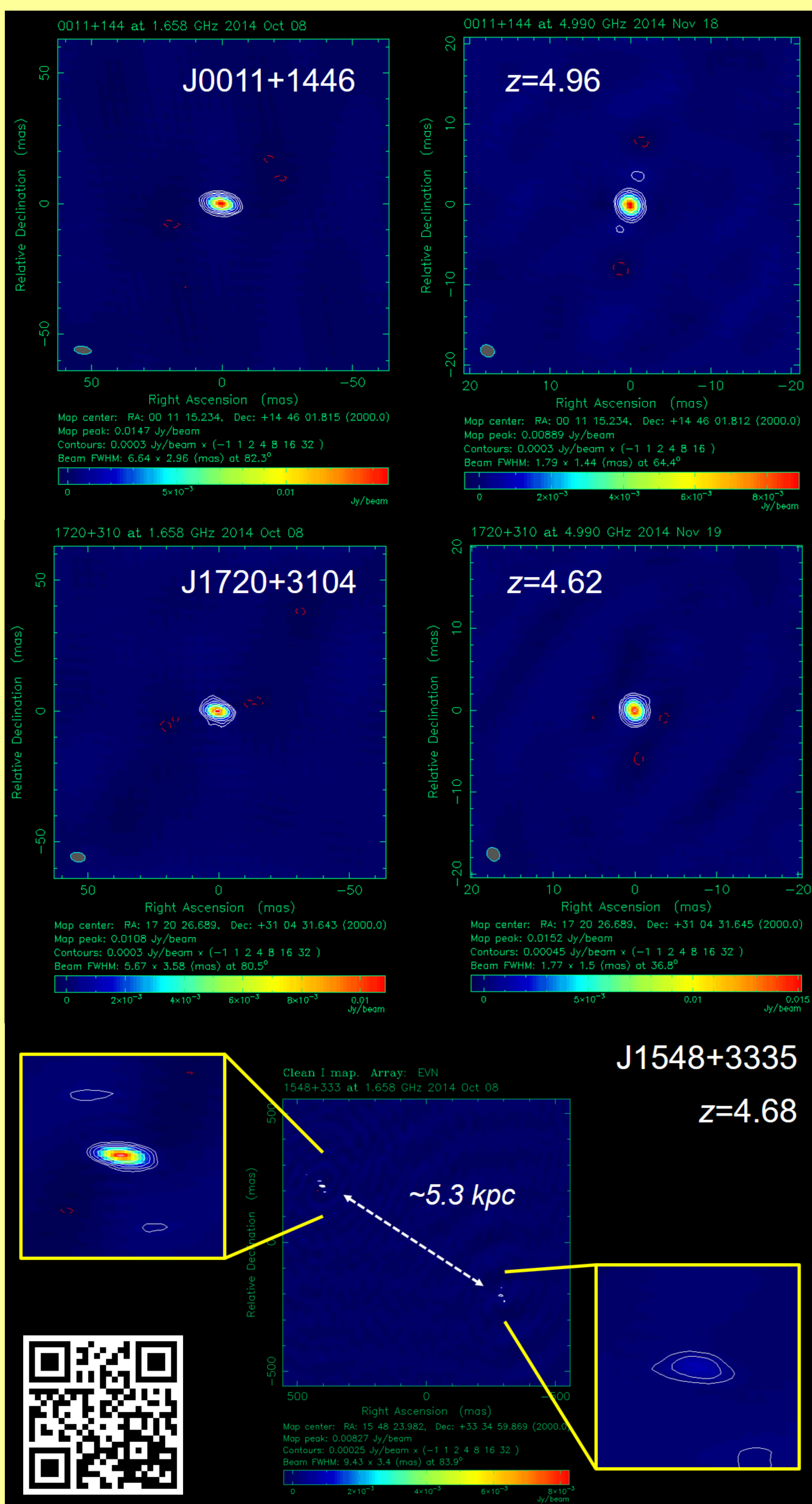
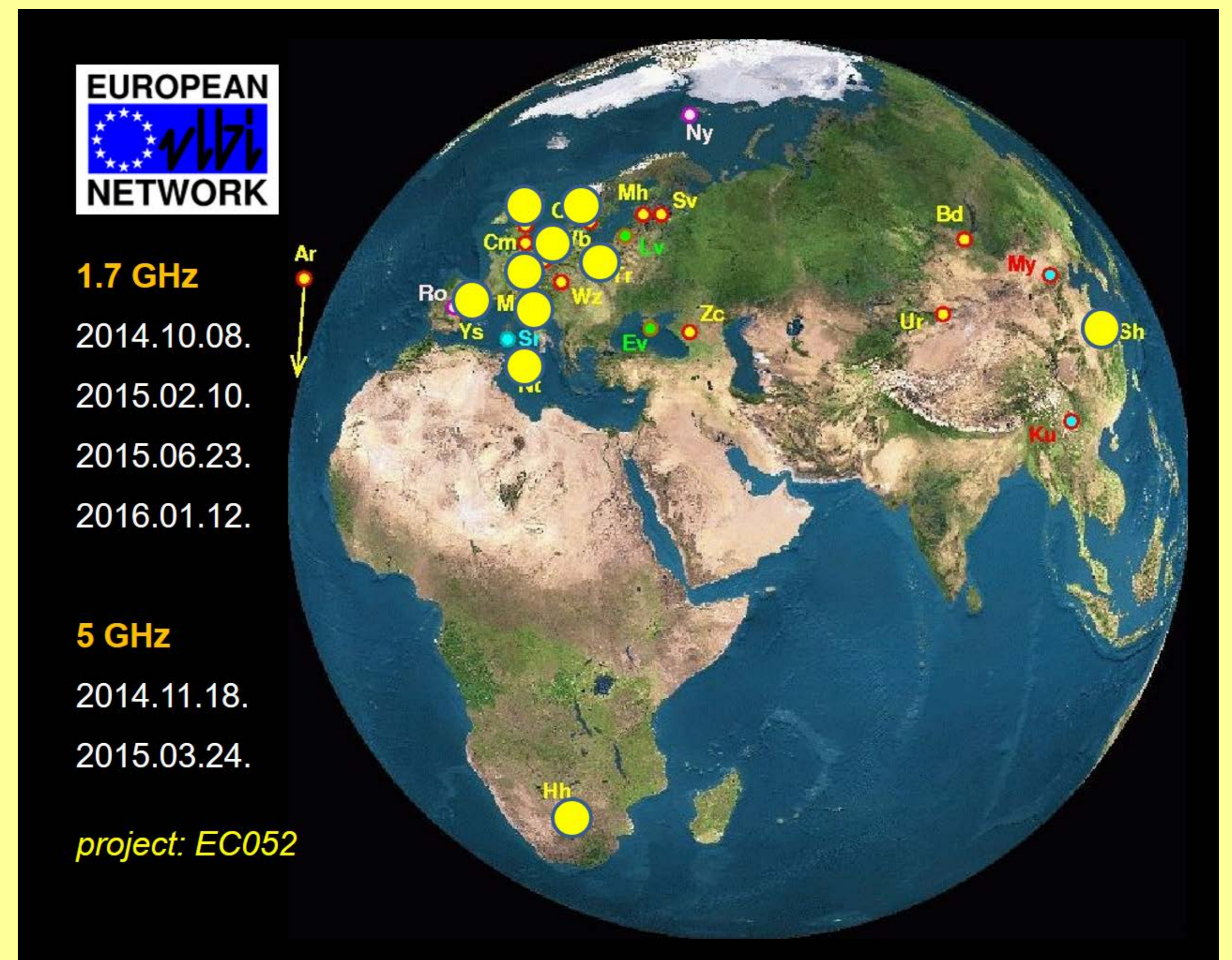
# Compact radio quasars at $z > 4.5$ observed with the EVN

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## Why are radio quasars at extremely high redshifts interesting?

- At redshifts above  $z > 4.5$ , the Universe was younger than  $\sim 10\%$  of its present age
- Active galactic nuclei (AGN) powered by supermassive black holes are seen at least until  $z \sim 7$
- High-redshift quasars are associated with the youngest supermassive black holes in the Universe
- Some of them are **radio-loud** and thus can be studied with VLBI at the highest possible resolution; radio provides an unobscured view of the nucleus
- They can be used to test the predictions of **cosmological models** (angular size-redshift and apparent proper motion-redshift relation)
- They can be used to set constraints on **black hole growth** in the early Universe and better understand **galaxy evolution**
- Recently we increased the number of VLBI-imaged AGN at  $z > 4.5$  by 50%, from 20 to 30, using e-EVN measurements



## Sample selection

- Optical quasars with known spectroscopic redshift  $z \geq 4.5$  (mostly from SDSS)
  - **Radio emission**, compact on arcsec scales (1.4-GHz flux density  $S > 5$  mJy from the VLA FIRST survey)
  - Not yet observed with VLBI
- ⇒ 10 new targets identified with  $4.5 \leq z \leq 5$
- Still very few similar objects are known*

## Results in a nutshell

- **All 10 targets are detected** with the EVN (9 at both frequencies, 1 at 5 GHz only)
- The images mostly show **featureless objects**, but the measured properties indicate a **mix of resolved and compact sources** on mas scales (*two examples are shown in the images*)
- The most intriguing one of the targets (J1548+3335) has a wide-separation **double structure** seen at 1.7 GHz (*see the image*) the brighter component that coincides with the optical quasar position is also detected at 5 GHz

## The sample of $z > 4.5$ radio sources observed with VLBI to date...

- ...consists of 30 objects and is **incomplete and heterogeneous**
- The highest redshift is  $z = 6.21$  (J1429+5447)
- The radio emission in one source (J1205-0742) originates from star formation; the others are AGN
- Based on **spectral index, brightness temperature** and (limited) **variability** information, roughly equal numbers of the rest are either **flat-spectrum radio quasars** or **steep-spectrum sources** (including two wide extended ones)
- Objects in the steep-spectrum category are most likely young sources with spectra peaking at GHz or  $\sim 100$  MHz frequencies

## Coming next:

- **Larger well-defined high-redshift radio quasar samples**
- **Better understanding of the selection effects**
- **Tackle the question of apparent overabundance of highly-beamed sources** (i.e. blazars) at very high redshifts

More details and further references:

Coppejans R. et al. 2016, *MNRAS*, in press (arXiv:1609.00575)